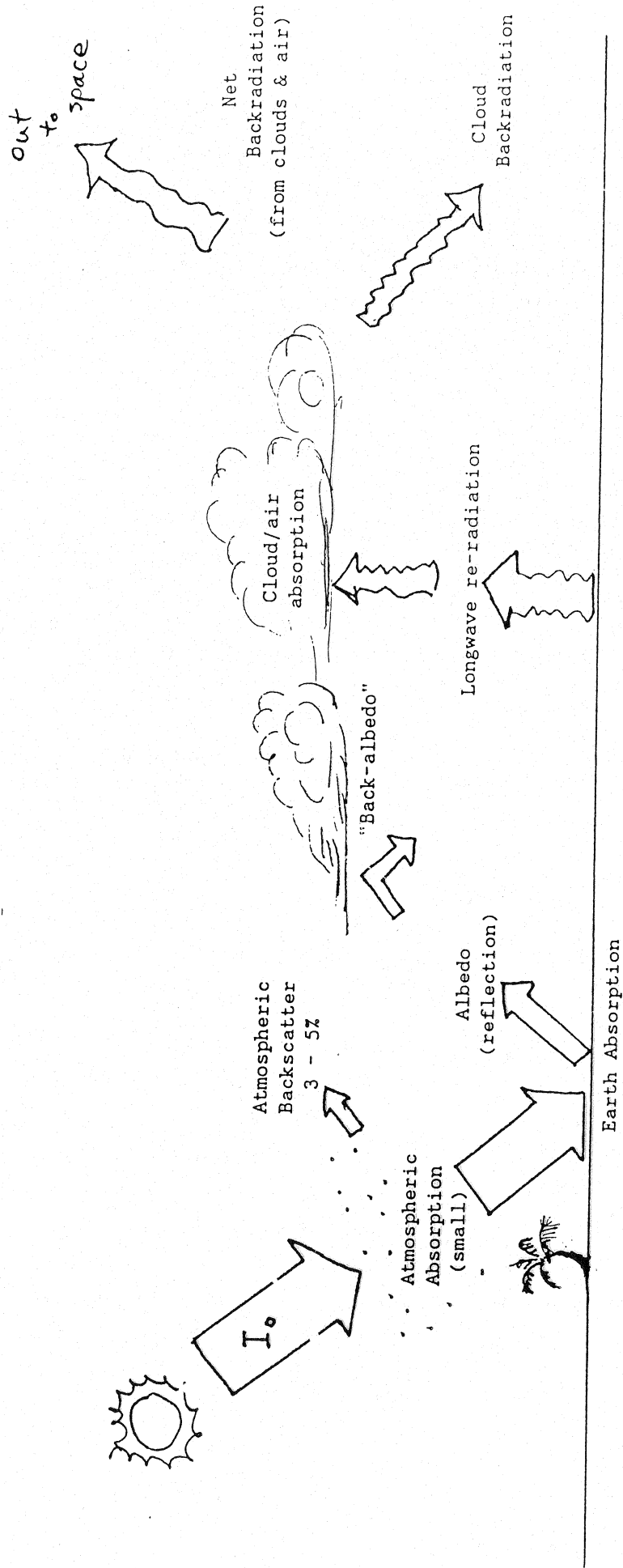


Energy Balance:

After light enters the lower atmosphere, it "bounces" around. ...

RADIATIVE ENERGY BALANCE IN THE LOWER ATMOSPHERE



A few important points:

# RADIATIVE FLUXES FOR A WATER BODY

SHORTWAVE = Visible light

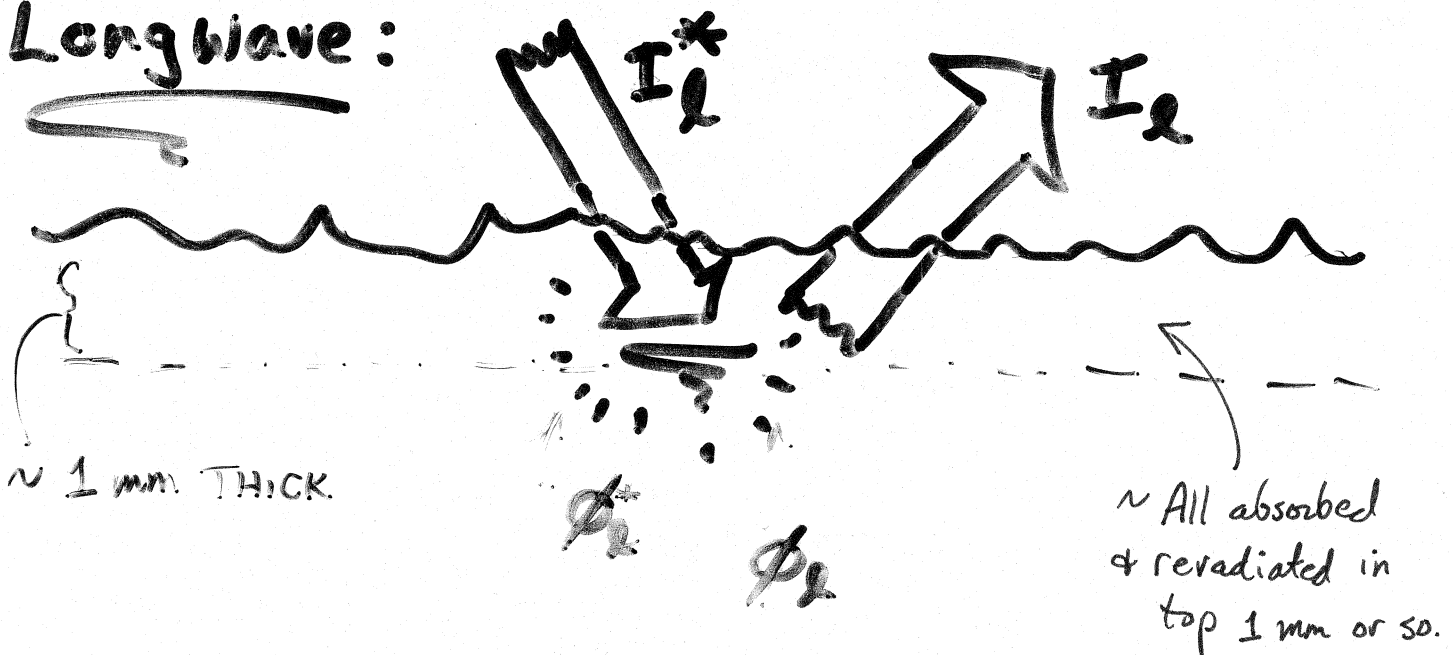
LONGWAVE = Infrared light

$$\phi_s = \text{Shortwave in} = I_s^*$$

$$\phi_l^* = \text{longwave in} = I_l^*$$

$$\phi_l = \text{longwave out} = I_l$$

Longwave:



$$\text{"OUT"} = (+) > 0$$

$$\text{"IN"} = (-) < 0$$

# TOTAL RADIATIVE INPUT

$$\Phi_r = \Phi_s + \Phi_l + \Phi_l^*$$

SHORT  
WAVE

LONG  
WAVE  
OUT

LONG  
WAVE  
IN

$$\approx -1000 \text{ W}\cdot\text{m}^{-2} \text{ (TYPICAL)}$$

SUNNY DAY, NOON

$$\approx +400 \text{ W/m}^2$$

WARM WATER, CLOUDLESS NIGHT

Be careful about "NET"

• NET RADIATION FLUX

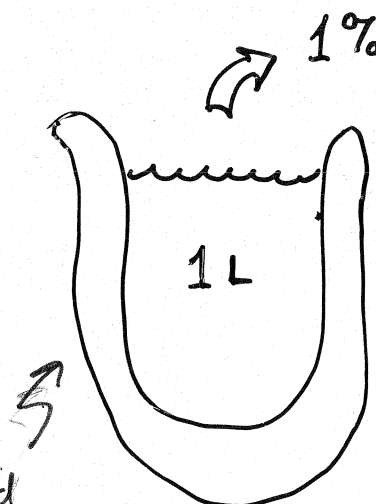
• NET HEAT FLUX

# EVAPORATIVE FLUX OF LATENT HEAT

$H_L$  = Latent heat of vaporization

For  $H_2O \approx 2,400 \text{ J/g}$   
(near  $20^\circ\text{C}$ )

E.g.:



Insulated Container  
"adiabatic"

Thermal Energy =  $24,000 \text{ J}$   
 $\approx 5,500 \text{ cal}$

$\Delta T = 5.5^\circ\text{C}$

$T_1 = 20^\circ\text{C} \rightarrow T_2 = 14.5^\circ\text{C}$

Or, to think another way

$24,000 \text{ J} =$  Energy of full tropical sunlight  
shining on the  $10 \text{ cm} \times 10 \text{ cm}$   
surface for  $\sim 40$  minutes

( $\circ\circ$  EVAP  $\approx$  SOLAR HEATING)

# EVAPORATION: Key, but can't measure directly

Relate heat loss to mass loss  
(gain) (gain)

$$\Phi_{\text{evap}} = \Phi_{\text{mass evap}} \cdot H_L$$

TYPICAL RANGE:

- 100  $\text{J m}^{-2} \text{s}^{-1}$  (W/m<sup>2</sup>) ~ MAX. CONDENSATION

+ 1000  $\text{W m}^{-2}$  ~ MAX. EVAPORATION

But, problems with measurement

Instead, use

Empirical relationships:

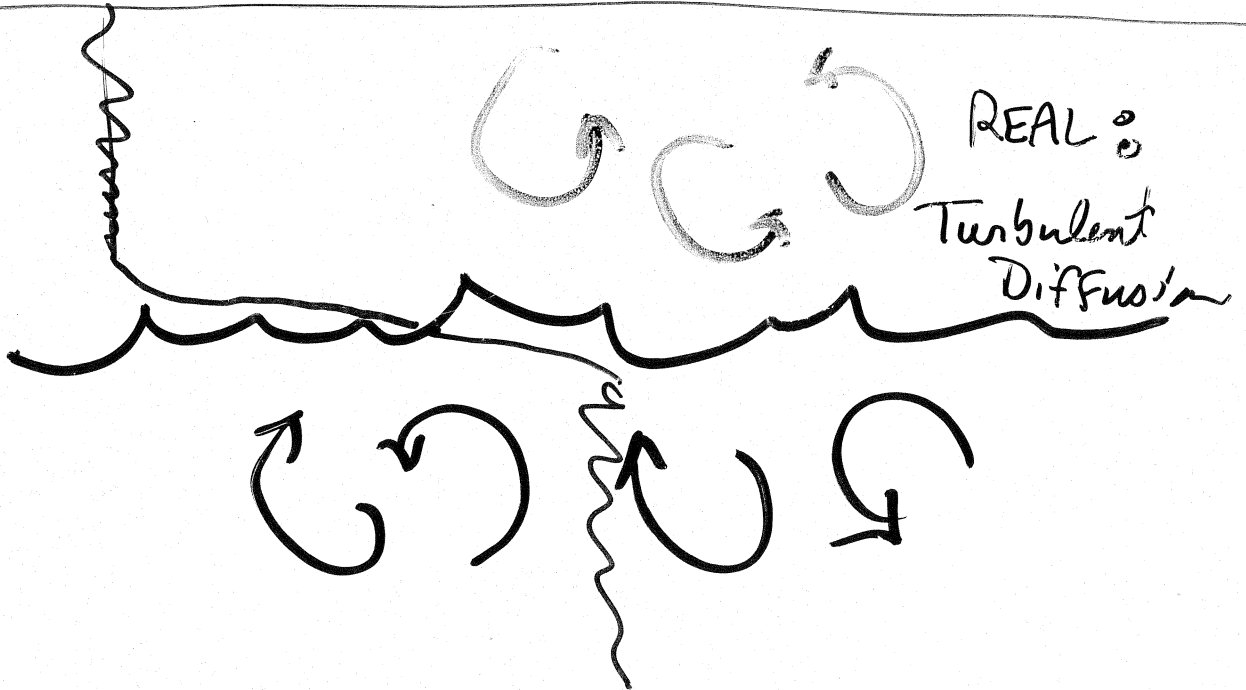
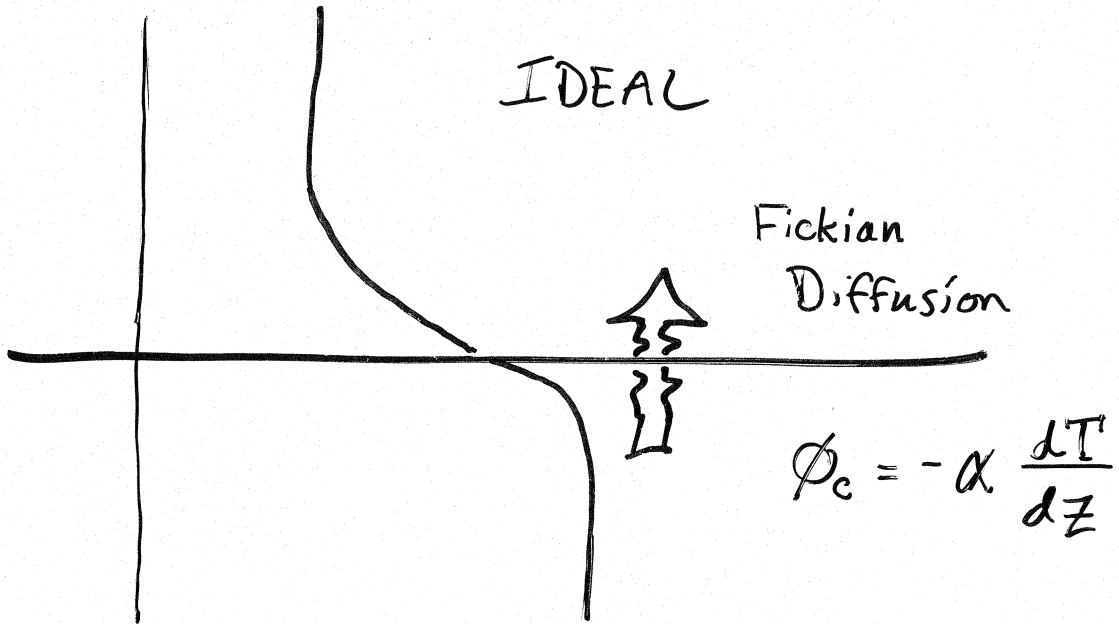
$$\Phi_e \approx H_L K_e (\rho_{vs} - \rho_{va})$$

$\rho_s$  = sat'd vapor density at  $T = T_{\text{surf}}$

$\rho_a$  = vapor density in air at ref. height

$K_e$  = empirical bulk transfer coeff.

# CONDUCTION



$$\phi_c = -K_H \frac{\Delta T}{\Delta z}$$

$\Delta T?$   $\Delta z?$

EMPIRICAL

$$\phi_c = \rho_a c_{pa} k_c (T_s - T_a)$$

# CONDUCTION (Sensible Heat Flux)

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## BOWEN RATIO:

$$\phi_c = R \phi_e$$

↑ CALCULATE/MEASURE  
 $\phi_e$   
Determine R.

$$R = C_p \left| \frac{T_s - T_z}{e_s - e_z} \right|$$

$$C_p = 61.0 \text{ Pa/}^\circ\text{C}$$

# NET NEAT FLUX

$$\Phi_{\text{net}} = \Phi_e + \Phi_c + \Phi_e$$

$$= \Phi_e + \Phi_c + \Phi_s^* + \Phi_e + \Phi_e^*$$

$$= \boxed{\Phi_e + \Phi_c + \Phi_e} + \boxed{\Phi_s^* + \Phi_e^*}$$

$T_s$  - Dependence

Astronomical



**FEEDBACK**



**"FORCING"**

$$\Phi_{\text{net}}[T_s] = \Phi_T(T_s) + \text{Constant}$$

